



## **Radiation-Induced transient effects in near Infrared focal plane arrays**

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## **Agenda**

- **Background/Problem**
- **Testing Goals and Strategy**
- **Test Data and Discussion**
- **Conclusions**


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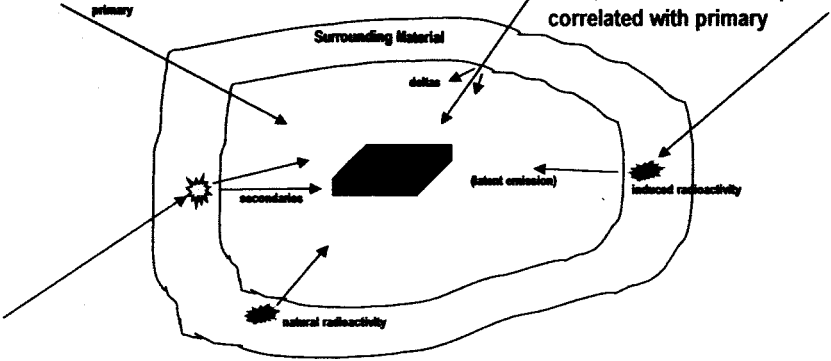


## Ionizing Particle Impacts to FPA



**+ Secondaries and delta electrons are time coincident with primary and have limited range**


**- But, not all deltas are spatially correlated with primary**




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## Even Small Transients can be Problematic for NIRCAM



- Read noise requirement is very low
- Essentially every primary particle and every secondary particle causes a transient that exceeds noise level
- Cosmic ray rejection algorithms can tolerate limited number of hits within integration time
- Problem is exacerbated by:
  - Crosstalk (charge spreading to neighboring pixels)
  - Multi-pixel hits (e.g., hit detector and ROIC in different pixels)
  - Secondary particles that are not spatially correlated to primary

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## TESTING GOALS, STRATEGY AND APPROACH

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## Transient Test Objectives

- **Characterize proton single events as function of energy and angle of incidence**
  - Pulse height distributions provide information for model calibration
- **Measure charge spread (crosstalk) to adjacent pixels**
  - Key parameter for determining number of disturbed pixels
- **Assess transient recovery time**
  - Look for long transients (collection of ionization-induced charge or persistence of radiation-induced dark current)
  - Characterize reset after hit

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## Devices Tested

- Test ROIC without detectors and test SCA (detectors plus ROIC) to separate effects
- Test 1024x1024 versions (H1RG and SB291)
  - Subset of identical circuitry on 2048x2048 versions

Device	Test Date	Energies	Angles	Comments
H1RG ROIC	5-16-02	30, 63	0	Single event and dose
H1RG SCA	10-16-02	30, 63	0, 45, 67	Single event, dose and secondaries from Al
SB291 ROIC	2-11-03	30, 63	0, 45, 67	Single event and dose
SB291 SCA	3-31-03	30, 63	0, 45, 67	Single event, dose and secondaries from Al

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## Transient Test Strategy

- Use 30 and 63 MeV protons
- Use 0, 45 and 67 degree incidence
- Use low flux for single events
  - 1e3 to 1e5 p/cm<sup>2</sup>-s range
- Use quilt-mode readout of 20x100 subarray
- Multiple samples at 10 Hz (100 ms) using variable integration time Fowler-mode integrations (reset, read<sub>1</sub>, read<sub>2</sub>, ... read<sub>i</sub>, reset)

Mode	#Reads	Time Between Reset (ms)
F1	2	200
F2	4	400
F5	10	1000
F10	20	2000

- Designed to capture transient recovery

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## 20\*x100 Subarray Used for Single Event Testing

\*Effectively 28x96 since  
includes 2 reference columns  
on each side

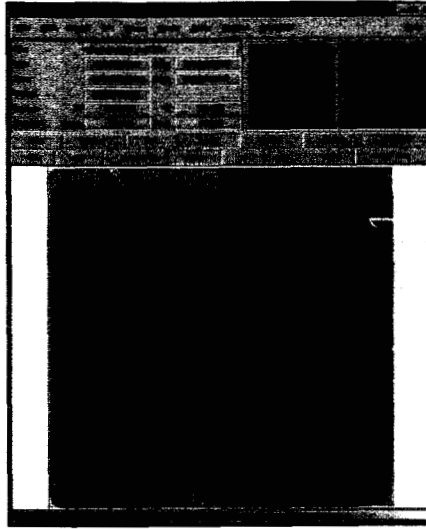


Figure 3. Example of a full frame pedestal (cross) road digitized 162x1624 image. Pixel (0,0) is located in the lower left hand corner. The green box represents the location of the subarray used in single-event testing.

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## TEST DATA AND DISCUSION

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## Typical Proton-Induced Pulses



H1RG SCA  
30 MeV

0 Degrees

5331	170	4	19	-17	2	8
100	-74	-55	52	-32	-15	92
20	20	-67	368	430	740	8121
-4	-4	507	28057	7562	14058	4831
-80	38	67	788	16245	1791	129
32	-136	87	2489	14295	24719	847
12	-48	584	7287	673	321	47

-79	25	35	-7	-34	0	23
-17	-40	-5	37	13	-27	0
-12	71	12	291	233	-100	26
-27	22	253	17941	16257	555	56
-10	-19	72	5522	5580	9946	3109
32	22	-53	96	245	6379	2290
-40	-123	14	-13	4	-180	-450

379	5	61	13	-107	-43	67
470	252	194	353	23	194	328
43	216	11306	26399	1941	14628	5209
14	-52	467	36916	2939	15379	5436
43	-15	12	488	-20	122	433
4178	1455	223	13403	2251	382	22117
12017	3043	62	85	-48	8	4288

67 Degrees

-34	-23	70	47	-64	-3	-23
23	44	-40	-20	7	17	-17
9	-1	382	15663	12690	448	29
17	-20	368	16398	14290	488	27
-27	26	36	249	215	48	-1
-3	0	23	0	-18	22	281
7	-10	37	17	-60	10	50

-64	-13	34	-34	0	44	-30
20	52	-20	-3	47	-37	-40
33	-6	197	415	72	-86	3
19	250	15232	28544	870	46	6
-4	56	4270	4773	284	19	50
-3	-30	40	57	10	-111	-821
7	44	-17	-17	-20	23	-23

-30	-3	64	-20	-13	3	-17
-64	-7	34	20	-23	-20	40
-44	-15	52	531	52	2	-13
7	42	2452	38147	4144	139	-25
-7	29	1505	17378	2880	115	31
10	-3	14	272	58	11	-40
13	37	-20	-37	20	17	-7

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## Typical Proton-Induced Pulses



SB291 SCA  
30 MeV

0 Degrees

-2	-2	6	-2	12	-2	-12
18	-27	6	-18	2	-2	10
12	2	2	-4	2	2	-16
20	-9	-16	413	344	-12	-5
0	0	0	145	86	8	4
-18	23	6	113	94	14	-8
-43	105	16	8652	826	-6	-17

-1	-4	-4	-6	8	-8	12
-1	14	0	-10	0	18	2
-1	2	-18	4	12	-14	12
-1	191	13488	16126	64	-18	123
-2	18	861	230	-1	-1	0
-1	-10	-4	-4	10	46	2
-1	14	133	2322	14	-8	-10

-10	-6	16	29	-18	-23	2
4	-22	23	-10	18	25	-41
-12	-13	507	732	32	-10	-10
4	11	4826	7094	39	-16	6
2	-4	-10	2	-6	-20	14
6	20	-23	-10	6	-2	-6
-8	6	4	-6	6	-8	-16

67 Degrees

20	-4	2	-27	10	23	-4
-20	0	4	2	18	22	-43
-10	4	2	12	-8	-8	4
1670	1690	11876	15894	77	15	135
2258	2620	1368	3275	30	-29	6
43	-37	4	-2	16	25	-16
-14	-27	22	12	-10	-4	0

8	14	-22	2	-8	-2	16
388	397	22	173	769	54	701
34	51	-13	-26	13	61	1254
1048	2338	393	22447	7556	226	109
1118	1392	229	14718	11197	62	-27
-16	-3	28	-145	-11	12	2
10	35	-23	16	6	-8	-27

6	14	-23	25	12	-6	-6
16	-10	-4	6	4	13	-6
701	1013	46	2089	1981	-3	-22
1254	2515	247	22516	12985	90	-5
109	275	46	-40	6	-6	-4
-27	180	111	-15	13	2	2
2	2	0	6	-6	-4	-14

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## Hit Size Expected to Scale With Proton Energy and Angle of Incidence



- $dE/dX$  { Linear Energy Transfer (LET) } is function of energy and material - lower at 63 MeV than 30 MeV
  - HgCdTe: 8.43 keV/um at 30 MeV; 4.91 keV/um at 63 MeV (x0.58)
  - InSb: 7.08 keV/um at 30 MeV; 4.09 Kev/um at 63 MeV (x0.58)
  - Si: 3.42 keV/um at 30 MeV; 1.92 keV/um at 63 MeV (x0.56)
- Charge generated  $\sim$  LET \* Path
- Path through device scales as  $\cos(\text{angle})$ 
  - x 1.41 at 45 degrees
  - x 2.56 at 67 degrees

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## PULSE HEIGHT DISTRIBUTIONS



- Difference between first read after reset and final read
- Data histogrammed into 100 e bins
- Analysis for two types of distributions:
  - Distribution of pixel charges
    - Some charge from hit pixels counted in neighboring pixels
  - Distribution of total hit charges
    - Hits identified and all charge from neighboring pixels added



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## SCA Pulse Height Distribution

Distribution of pixel charges

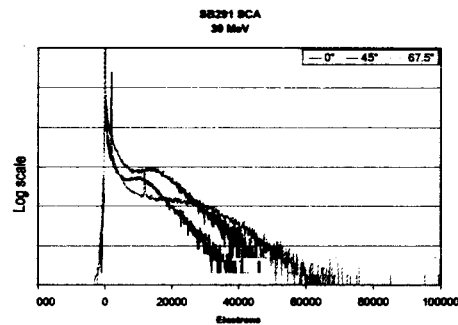
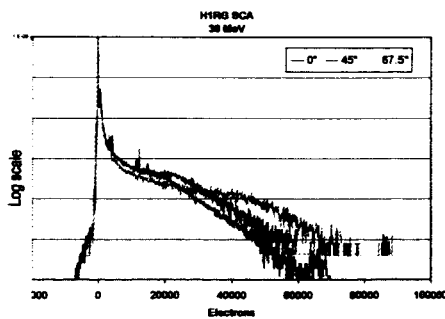
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## SCA Comparison at Various Angles 30 MeV



- SCA pulses are larger than ROIC pulses
- Pulses scale with angle of incidence
- H1RG pulses are larger than SB291 pulses

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## TOTAL CHARGE DISTRIBUTIONS

Total Charge : Charge to hit pixel and  
affected neighbors summed

Peak Charge: Charge to hit pixel only

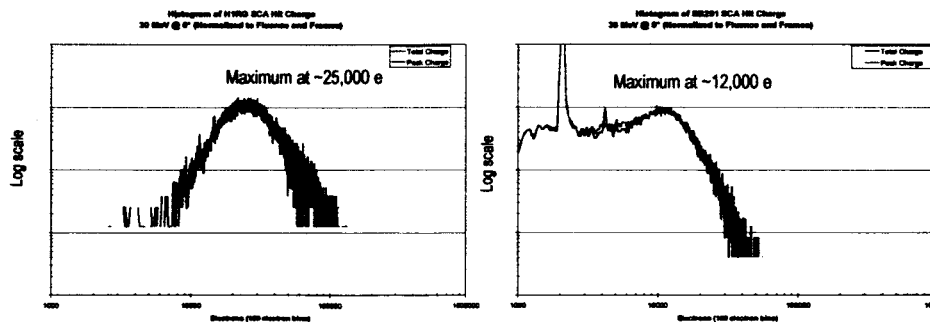
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## SCA Total Charge Distributions 30 MeV, 0 Degrees



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## Total Charge Peaks Scale With Angle and Energy as Expected



Peak Locations				Angle Scaling			Energy Scaling		
Energy	Angle	H1RG	H1RG	Expected	H1RG	H1RG	Expected	H1RG	H1RG
30	0	23000	23000	1.00	1.00	1.00	1.00	1.00	1.00
30	45	25000	25000	1.41	1.38	1.09	1.00	1.00	1.00
30	67	48000	48000	2.56	2.38	2.09	1.00	1.00	1.00
63	0	9000	9000	1.00	1.00	1.00	0.58	0.67	0.39
63	45	13000	13000	1.41	1.25	1.44	0.58	0.63	0.52
63	67	25000	25000	2.56	2.38	2.78	0.58	0.63	0.52

H1RG SCA pulses are somewhat larger than SB291SCA pulses

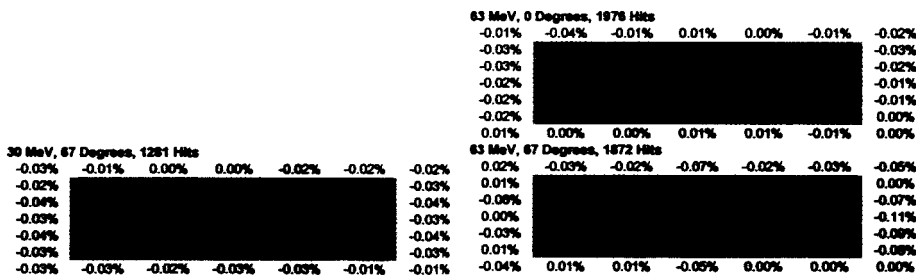
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## Measured Proton Crosstalk in H1RG SCA



- Hits randomly distributed across pixel but all at same angle
- Hits are stacked by registration to hit pixel, not to hit centroid
- In some cases, charge is still above noise even at 2 pixels out from hit pixel

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## Measured Proton Crosstalk in SB291 SCA



30 MeV, 0 Degrees, 3755 Hits



63 MeV, 0 Degrees, 1385 Hits



63 MeV, 67 Degrees, 814 Hits



- Hits randomly distributed across pixel but all at same angle
- Hits are stacked by registration to hit pixel, not to hit centroid
- In some cases, charge is still above noise even at 2 pixels out from hit pixel

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## Observations



- Pulses generally scale with energy and with angle as expected
- Unipolar pulses in SB291 (same polarity as detector)
- Bipolar polarity pulses in H1RG ROIC
- Pulses are smaller in ROIC than SCA for both H1RG and SB291
- Pulses with same polarity as detector have comparable size for H1RG ROIC and SB291 ROIC
- Pulses are larger in H1RG SCA than SB291 SCA
- Crosstalk is larger in SB291 ROIC than H1RG
- Crosstalk is larger in H1RG SCA than SB291
- Hit pixel recovery <100 ms or upon reset

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## General Conclusions



- **Whatever technology chosen, JWST will have to live with cosmic ray hits**
- **Overall transient responses are similar at SCA level**
  - ROIC hits are larger for SB291 than H1RG
  - H1RG SCA hits are larger (apparently due to detector)
  - H1RG proton crosstalk is worse (probably related to smaller pitch)
- **Note that smaller pixels would have lower hit probability in space environment but more crosstalk**